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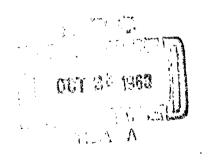
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DERIVATION OF EQUATIONS FOR CONVERTING FROM GEODETIC COORDINATES TO GEOCENTRIC COORDINATES

by F. T. Heuring



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Derivation Of Equations
For Converting
From Geodetic Coordinates
To Geocentric Coordinates

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THE JOHNS HOPKINS UNIVERSITY
APPLIED PHYSICS LABORATORY

H21 GEORGIA AVENUE SILVER SPRING, MARYLAND

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DERIVATION OF EQUATIONS FOR CONVERTING FROM GEODETIC COORDINATES TO GEOCENTRIC COORDINATES

F. T. Heuring

In the A.P.L. orbit computation programs, the TRANET Tracking Stations are specified in a geocentric coordinate system, whereas, particular positions (such as a TRANET Tracking Site) over the Earth are expressed initially in a geodetic coordinate system. In order to acquire geocentric coordinates from a given set of geodetic coordinates a set of transformation equations were derived.

Section I will define the notation, and Section II will embody the derivation of the transformation equations.

I. Notation*

Let:

 ϕ_{G_3} = geodetic latitude of i-th tracking site in its local datum,

 $\lambda_{G_{\hat{i}}}$ = geodetic longitude of i-th tracking site in its local datum,

h, = elevation of i-th tracking site above (below) geoid,

 H_i = geoidal height of i-th tracking site in its local datum,

 ξ_i = deflection in meridian at i-th tracking site,

 η_i = deflection in prime vertical at i-th tracking site,

a = equatorial radius of the local datum spheroid of the i-th
 tracking site, scaled by R_a,

b_i = polar radius of the local datum spheroid on the i-th tracking site, scaled by R_o,

^{*}See References 1 and 2 for definition of geodetic, datum, etc.

 x_{G_1} , y_{G_1} , z_{G_2} = cartesian coordinates, scaled by R_O , on i-th datum spheroid as specified by tracking site ϕ_G and λ_G , (cartesian origin identical to i-th datum origin),

x_{H_i}, y_{H_i}, z_{H_i} = cartesian coordinates, scaled by R_o, on geoid as specified by tracking site H, (cartesian origin identical to i-th datum origin),

x_E, y_E, z_E = cartesian coordinates, scaled by R_O, of tracking site on earth's surface, (cartesian origin identical to i-th datum origin),

 Δx_i , Δy_i , Δz_i = center of spheroid of the i-th tracking site datum in the A.P.L. Datum, scaled by R_o ,

$$\zeta_{G_i} = (x_{G_i}^2 + y_{G_i}^2)^{\frac{1}{2}}$$

 R_{\odot} = equatorial radius of A.P.L. Datum spheroid,

x_{ci}, y_{ci}, z_{ci} = cartesian coordinates of tracking site in A.P.L. geocentric coordinates, scaled by R_O,

 r_{ci} = radius of i-th tracking site in A.P.L. geocentric coordinates, scaled by R_{ci} ,

 ϕ_{c} = latitude of i-th tracking site in A.P.L. geocentric coordinates,

 λ_{c}^{1} = longitude of i-th tracking site in A.P.L. geocentric coordinates,

 $\zeta_{c_{i}}^{1} = (x_{c_{i}}^{2} + y_{c_{i}}^{2})^{\frac{1}{2}}.$

II. Derivation

A. Given ϕ_{G_i} , λ_{G_i} , a and b, conversion to x_{G_i} , y_{G_i} , z_{G_i} and ζ_{G_i} is as follows. Using the equation for an ellipse

$$\frac{\zeta_{G_{1}}^{2}}{a_{1}^{2}} + \frac{z_{G_{1}}^{2}}{b_{1}^{2}} = 1,$$

in particular the ellipse is a meridional plane of the i-th datum; differentiate z_{G_i} with respect to ζ_{G_i}

$$\frac{\partial^{z}G_{i}}{\partial\zeta_{G_{i}}} = -\frac{b_{i}^{2}}{a_{i}^{2}} \frac{\zeta_{G_{i}}}{z_{G_{i}}}.$$

But (see Figure 1A),

$$\frac{\partial^{z}G_{i}}{\partial \zeta_{G_{i}}} = -\frac{1}{\tan \varphi_{G_{i}}}$$

from which by algebraic manipulation (Figure 1B),

$$\zeta_{G_{i}} = \frac{a_{i}}{(1 + (\frac{b_{i}}{a_{i}})^{2} \tan^{2} \varphi_{G_{i}})^{\frac{1}{2}}},$$

after which,

$$\mathbf{x}_{G_{\underline{i}}} = \zeta_{G_{\underline{i}}} \cos \lambda_{G_{\underline{i}}}$$

$$\mathbf{y}_{G_{\underline{i}}} = \zeta_{G_{\underline{i}}} \sin \lambda_{G_{\underline{i}}}$$

$$\mathbf{z}_{G_{\underline{i}}} = \zeta_{G_{\underline{i}}} \frac{b_{\underline{i}}^{2}}{a_{\underline{i}}^{2}} \tan \phi_{G_{\underline{i}}}.$$

$$(1)$$

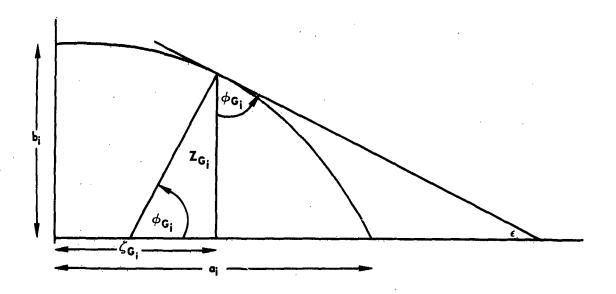


Figure 1A Meridian Plane in i-th Datum

$$\frac{1}{\tan \phi_{G_i}} = \cot \phi_{G_i} = \tan \epsilon - \frac{\partial z_{G_i}}{\partial \zeta_{G_i}}$$

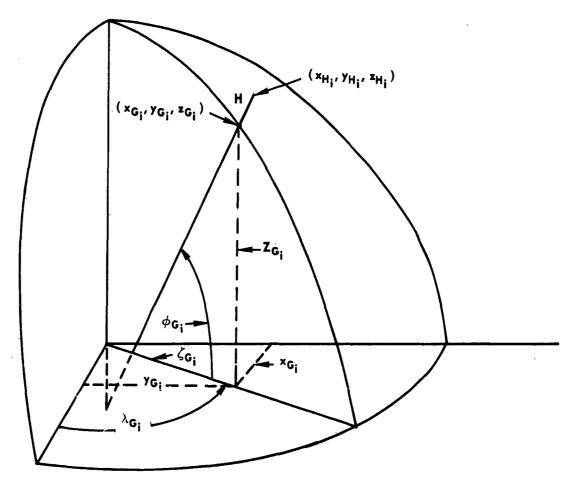


Figure 1B Pictorial view of geodetic (ϕ_{G_i} , λ_{G_i}), cartesian "geodetic" (\mathbf{x}_{G_i} , \mathbf{y}_{G_i} , \mathbf{z}_{G_i}) and cartesian "geoidal" (\mathbf{x}_{H_i} , \mathbf{y}_{H_i} , \mathbf{z}_{H_i}) coordinates.

B. Compute $x_{H_{\dot{1}}}$, $y_{H_{\dot{1}}}$, $z_{H_{\dot{1}}}$ (Figure 1B). $H_{\dot{1}}$ is an extension of the normal to the spheroid, consequently,

$$\begin{aligned} \mathbf{x}_{\mathbf{H}_{\mathbf{i}}} &= \mathbf{x}_{\mathbf{G}_{\mathbf{i}}} + \mathbf{H}_{\mathbf{i}} \cos \phi_{\mathbf{G}_{\mathbf{i}}} \cos \lambda_{\mathbf{G}_{\mathbf{i}}} \\ \mathbf{y}_{\mathbf{H}_{\mathbf{i}}} &= \mathbf{y}_{\mathbf{G}_{\mathbf{i}}} + \mathbf{H}_{\mathbf{i}} \cos \phi_{\mathbf{G}_{\mathbf{i}}} \sin \lambda_{\mathbf{G}_{\mathbf{i}}} \\ \mathbf{z}_{\mathbf{H}_{\mathbf{i}}} &= \mathbf{z}_{\mathbf{G}_{\mathbf{i}}} + \mathbf{H}_{\mathbf{i}} \sin \phi_{\mathbf{G}_{\mathbf{i}}}. \end{aligned} \tag{2}$$

C. Compute x_{E_i} , y_{E_i} , z_{E_i} by considering h_i , ξ_i and η_i (see Figure 21).

$$\begin{aligned} & \mathbf{x_{E_{i}}} &= \mathbf{x_{H_{i}}} + \mathbf{h_{i}} \cos (\phi_{G_{i}} + \xi_{i}) \cos (\lambda_{G_{i}} + \Delta \lambda_{i}) \\ & \mathbf{y_{E_{i}}} &= \mathbf{y_{H_{i}}} + \mathbf{h_{i}} \cos (\phi_{G_{i}} + \xi_{i}) \sin (\lambda_{G_{i}} + \Delta \lambda_{i}) \\ & \mathbf{z_{E_{i}}} &= \mathbf{z_{H_{i}}} + \mathbf{h_{i}} \sin (\phi_{G_{i}} + \xi_{i}). \end{aligned} \tag{3}$$

From law of cosines for spherical triangles (Figure 12.), $\Delta\lambda_{\, \rm i}$ can be approximated.

$$\cos \Delta \lambda_{i} = \frac{\cos \eta_{i} - \sin^{2} (\varphi_{G_{i}} + \xi_{i})}{\cos^{2} (\varphi_{G_{i}} + \xi_{i})}. \tag{4}$$

(Restrict $\Delta \lambda_i$ to have the same sign as η_i).

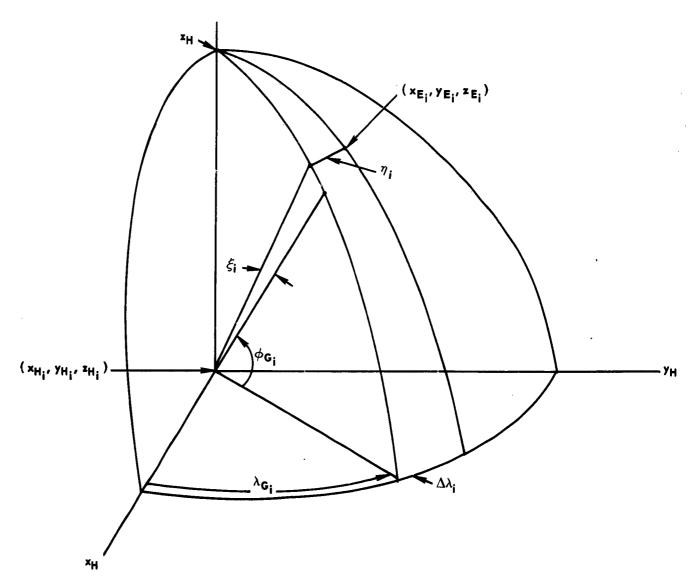


Figure 2 Diagram of the deflections of the vertical (ξ_i and η_i) and the associated quantities necessary to acquire the cartesian coordinates on the geoid from earth surface cartesian coordinates.

D. Let us simplify by expanding small quantities. Assume:

$$\xi_{i}, \eta_{i} \leq 30"^{*} \text{ (of arc)};$$

and

$$1^{\circ} < |\phi_{G_{i}}| < 89^{\circ};$$

and only take quantities of magnitude ξ_i , η_i and $\Delta\lambda_i$ to second order.

$$\cos \xi_{i} \doteq 1 - \frac{\xi_{i}^{2}}{2}, \sin \xi_{i} \doteq \xi_{i}$$

$$\cos \eta_{i} \doteq 1 - \frac{\eta_{i}^{2}}{2}, \sin \eta_{i} \doteq \eta_{i}$$

$$\cos \Delta \lambda_{i} \doteq 1 - \frac{\Delta \lambda_{i}^{2}}{2}$$

thus.

$$\cos^{2}(\varphi_{G_{i}} + \xi_{i}) = \left[\cos\varphi_{G_{i}} \left(1 - \frac{\xi_{i}^{2}}{2}\right) - \xi_{i}\sin\varphi_{G_{i}}\right]^{2}$$

$$= \left(1 - \frac{\xi_{i}^{2}}{2}\right)^{2}\cos^{2}\varphi_{G_{i}} + \xi_{i}^{2}\sin^{2}\varphi_{G_{i}}$$

$$- 2 \xi_{i}\left(1 - \frac{\xi_{i}^{2}}{2}\right)\sin\varphi_{G_{i}}\cos\varphi_{G_{i}}$$

$$= \cos^{2}\varphi_{G_{i}} - \xi_{i}\sin^{2}\varphi_{G_{i}} - \xi_{i}^{2}\cos^{2}\varphi_{G_{i}} + 3rd \text{ order}$$
 (5)

From a personal communication with Mr. L. Simmons, U.S.C. and G.S., deflection of 30" exist but are in general uncommon.

(9)

+ 3rd order

 $+ \xi_1^2 \cos 2 \phi_{G_1}^4$

sin 2 $\phi_{\mathtt{G}_{\mathtt{l}}}$

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= sin² a

 $\cos \phi_{\mathbf{G_1}}$

+2**%** sin ϕ_{i} c

og.

+ 52 cos c

∸) sin² ه_G +

(1)

cos $\phi_{G_{\underline{1}}}$

S_i sin ϕ_{G_1}

 $= \cos \phi_{G_1}$

S_i sin ϕ_{G} :

 $\cos \left(\phi_{\mathcal{G}_{1}} + \xi_{1} \right) = \cos \phi_{\mathcal{G}_{1}}$

(8)

 $+ \xi_1 \cos \varphi_{G_1}$

= sin $\phi_{\mathrm{G}_{\mathtt{i}}}^{}$

+ 5₁ cos φ_{G1} = $\sin(\phi_{G_1} + \xi_1) = \sin\phi_{G_1}$

= $\sin^2 (\varphi_{G_1} + \xi_1) + \cos^2 (\varphi_{G_1} + \xi_1) [1 -$

9

 $^{\circ} + \xi_{i} \cos \varphi_{G_{i}}$

 $\sin^2(\varphi_{G_1} + \xi_1) = [\sin\varphi_{G_1} (1 - -$

from equation (4):

and using equation (7),

$$\Delta \lambda_{1} = \frac{\eta_{1}}{\cos^{2} \phi_{G_{1}}} \left[\frac{1}{1 - \xi_{1} \tan^{2} \phi_{G_{1}}} - \frac{1}{\xi_{1}} \right] = \frac{\eta_{1}}{\cos^{2} \phi_{G_{1}}} \left[\frac{1 + \xi_{1} \tan^{2} \phi_{G_{1}}}{1 + \xi_{1} \tan^{2} \phi_{G_{1}}} + \frac{\xi_{2}^{2}}{2} + \frac{\xi_{1}^{2}}{2} \tan^{2} \phi_{G_{1}} \right]$$

$$= \eta_{1} \sec^{2} \phi_{G_{1}} \left[\frac{1 + \xi_{1} \tan^{2} \phi_{G_{1}}}{1 + \xi_{1} \tan^{2} \phi_{G_{1}}} + \frac{1}{3} \tan^{2} \phi_{G_{1}} \right] + 3rd \text{ order.}$$

6)

Further,

$$\sin \left(\lambda_{\mathbf{G_1}} + \Delta \lambda_{\mathbf{1}} \right) = \sin \lambda_{\mathbf{G_1}} \left(1 - \frac{\eta_{\mathbf{1}}^2 \sec^2 \phi_{\mathbf{G_1}}}{2} \right) + \cos \phi_{\mathbf{G_1}} \eta_{\mathbf{1}} \sec \phi_{\mathbf{G_1}} \left(1 + \xi_{\mathbf{1}} \tan \phi_{\mathbf{G_1}} \right)$$

$$= \sin \lambda_{\mathbf{G_1}} + \eta_{\mathbf{1}} \frac{\cos^3 \lambda_{\mathbf{G_1}}}{\cos^3 \phi_{\mathbf{G_1}}} + \frac{\eta_{\mathbf{1}}}{\cos^3 \phi_{\mathbf{G_1}}} \left[\xi_{\mathbf{1}} \cos \lambda_{\mathbf{G_1}} \sin \phi_{\mathbf{G_1}} - \frac{\eta_{\mathbf{1}}}{2} \sin \lambda_{\mathbf{G_1}} \right] + 3rd \text{ order}$$

(10)

$$\cos \left(\lambda_{G_{\underline{1}}} + \Delta \lambda_{\underline{1}} \right) = \cos \lambda_{G_{\underline{1}}} - \sin \lambda_{G_{\underline{1}}} \sec \phi_{G_{\underline{1}}} \left(1 + \xi_{\underline{1}} \tan \phi_{G_{\underline{1}}} \right) \eta_{\underline{1}} - \frac{\eta_{\underline{1}}^2}{2} \sec^2 \phi_{G_{\underline{1}}} \cos \lambda_{G_{\underline{1}}}$$

$$=\cos\lambda_{\mathbf{G_{1}}}-\eta_{1}\frac{\sin\lambda_{\mathbf{G_{1}}}}{\cos\phi_{\mathbf{G_{1}}}}-\frac{\eta_{1}}{\cos^{2}\phi_{\mathbf{G_{1}}}}\left[\sum_{\mathbf{S_{1}}\sin\lambda_{\mathbf{G_{1}}}\sin\phi_{\mathbf{G_{1}}}}^{\mathrm{gin}}\phi_{\mathbf{G_{1}}}+\frac{\eta_{1}}{2}\cos\lambda_{\mathbf{G_{1}}}\right]+3\mathrm{rd\ order.}$$

Using equations (1), (2), (3), (7), (8), (9), (10), and (11), $\mathbf{x_{E}}$, $\mathbf{y_{E}}$ and $\mathbf{z_{E}}$ can be expressed functions of the geodetic inputs $(\phi_{G_{1}}, \lambda_{G_{1}}, \mu_{1}, \mu_{1}, \eta_{1}, \xi_{1}, a_{1}, and b_{1})$

ω

 $\sin \lambda_{G_{i}}$ $= \zeta_{\mathbf{G_1}} \cos \lambda_{\mathbf{G_1}} + H_1 \cos \phi_{\mathbf{G_1}} \cos \lambda_{\mathbf{G_1}} + h_1 \left(\cos \phi_{\mathbf{G_1}} - \xi_1 \sin \phi_{\mathbf{G_1}}\right) \left(\cos \lambda_{\mathbf{G_1}} - \eta_1 \frac{1}{\cos \phi_{\mathbf{G_1}}}\right)$ $= \zeta_{\mathbf{G}_{\underline{1}}} \cos \lambda_{\mathbf{G}_{\underline{1}}} + (\mathbf{H}_{\underline{1}} + \mathbf{h}_{\underline{1}}) \cos \phi_{\mathbf{G}_{\underline{1}}} \cos \lambda_{\mathbf{G}_{\underline{1}}} - \mathbf{h}_{\underline{1}} (\xi_{\underline{1}} \sin \phi_{\mathbf{G}_{\underline{1}}} \cos \lambda_{\mathbf{G}_{\underline{1}}} + \eta_{\underline{1}} \sin \lambda_{\mathbf{G}_{\underline{1}}}) + 3rd \text{ order.}$

 $\mathbf{y_{E_1}} = \zeta_{\mathbf{G_1}} \sin \lambda_{\mathbf{G_1}} + \mathbf{H_1} \cos \phi_{\mathbf{G_1}} \sin \lambda_{\mathbf{G_1}} + \mathbf{h_1} (\cos \phi_{\mathbf{G_1}} - \xi_1 \sin \phi_{\mathbf{G_1}}) \left(\sin \lambda_{\mathbf{G_1}} + \eta_1 \frac{1}{\cos \phi_{\mathbf{G_1}}} \right)$

 $= \zeta_{\mathbf{d_1}} \sin \lambda_{\mathbf{d_1}} + (\mathbf{H_1} + \mathbf{h_1}) \cos \varphi_{\mathbf{d_1}} \sin \lambda_{\mathbf{d_1}} - \mathbf{h_1}(\xi_1 \sin \varphi_{\mathbf{d_1}} \sin \lambda_{\mathbf{d_1}} - \eta_1 \cos \lambda_{\mathbf{d_1}}) + 3rd \text{ order.}$

 $\frac{b_i^2}{a_i^2}$ tan ϕ_{G_i} + H_i sin ϕ_{G_i} + h_i (sin ϕ_{G_i} + ξ_i cos ϕ_{G_i}

+ 3rd order. $\frac{1}{2}$ tan ϕ_{G_1} + $(H_1 + h_1)$ sin ϕ_{G_1} + h_1 $\xi_1 \cos \phi_{G_1}$ F. The cartesian coordinates in the A.P.L. geocentric system are:

$$x_{c_{i}} = x_{E_{i}} + \Delta x_{i}$$

$$y_{c_{i}} = y_{E_{i}} + \Delta y_{i}$$

$$z_{c_{i}} = z_{E_{i}} + \Delta z_{i}$$

where Δx_i , Δy_i , and Δz_i are of second order, at best.

H. The cylindrical coordinates (z , ζ_c , λ_c) in the A.P.L. geocentric system are:

$$z_{c_{i}} = \zeta_{G_{i}} \frac{b_{i}^{2}}{2} \tan \varphi_{G_{i}} + (H_{i} + h_{i}) \sin \varphi_{G_{i}} + h_{i} \xi_{i} \cos \varphi_{G_{i}} + \Delta z_{i} + 3rd \text{ order (13)}$$

$$\zeta_{c_{i}}^{2} = x_{c_{i}}^{2} + y_{c_{i}}^{2}$$

After some algebraic manipulation and using the binominal expansion

$$\zeta_{c_{\underline{i}}} = \zeta_{G_{\underline{i}}} + (H_{\underline{i}} + h_{\underline{i}}) \cos \varphi_{G_{\underline{i}}} + \Delta x_{\underline{i}} \cos \lambda_{G_{\underline{i}}} + \Delta y_{\underline{i}} \sin \lambda_{G_{\underline{i}}} - h_{\underline{i}} \xi_{\underline{i}} \sin \varphi_{G_{\underline{i}}} \\
+ (H_{\underline{i}} + h_{\underline{i}}) \cos \varphi_{G_{\underline{i}}} \cdot \frac{1}{\zeta_{G_{\underline{i}}}} (\Delta x_{\underline{i}} \cos \lambda_{G_{\underline{i}}} + \Delta y_{\underline{i}} \sin \lambda_{G_{\underline{i}}}) \tag{14}$$

+ 3rd order.

In the derivation of $\lambda_{G_{\hat{1}}}$, no previously derived quantities were used as was the case with $\zeta_{c_{\hat{1}}}$. From Figure 3A, $h_{\hat{1}}$ is considered to be zero, thus the angle c can be approximated as follows:

$$\mathbf{e}_{1} + \mathbf{e}_{2} = \Delta \mathbf{x}_{i} \sin \lambda_{\mathbf{G}_{i}}$$

$$\mathbf{e}_{2} \doteq \Delta \mathbf{y}_{i} \cos \lambda_{\mathbf{G}_{i}} \qquad \text{where } \mathbf{e}_{1} \text{ and } \mathbf{e}_{2} \text{ are normal to } \zeta_{\mathbf{G}_{i}}, \text{ and }$$

$$\mathbf{e}_{1} \doteq \Delta \mathbf{x}_{i} \sin \lambda_{\mathbf{G}_{i}} - \Delta \mathbf{y}_{i} \cos \lambda_{\mathbf{G}_{i}}.$$

Since ϵ_1 considered, at best, second order,

$$\sigma = \frac{\epsilon_1}{\zeta_{G_1}}$$

and from the geometry,

$$\lambda_{c_{i}} = \lambda_{G_{i}} - \sigma = \lambda_{G_{i}} - \frac{1}{\zeta_{G_{i}}} (\Delta x_{i} \sin \lambda_{G_{i}} - \Delta y_{i} \cos \lambda_{G_{i}})$$
 (15)

Upon including the station elevation (h $_{\rm i})$ and deflection in the prime vertical ($\eta_{\rm i})$ (see Figure 3B)

$$\tau = h_i \sin \eta_i = h_i \eta_i$$
 (η_i is of magnitude + 30" of arc)

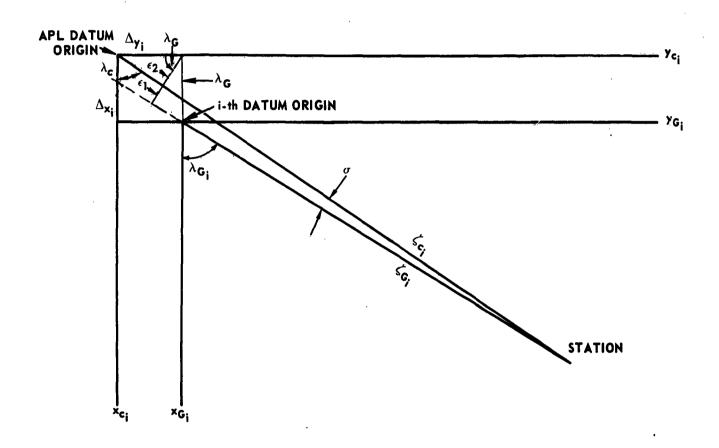


Figure 3A Diagram showing means of determining $\lambda_{\boldsymbol{c}}$ when $h_{\boldsymbol{i}}=0.$

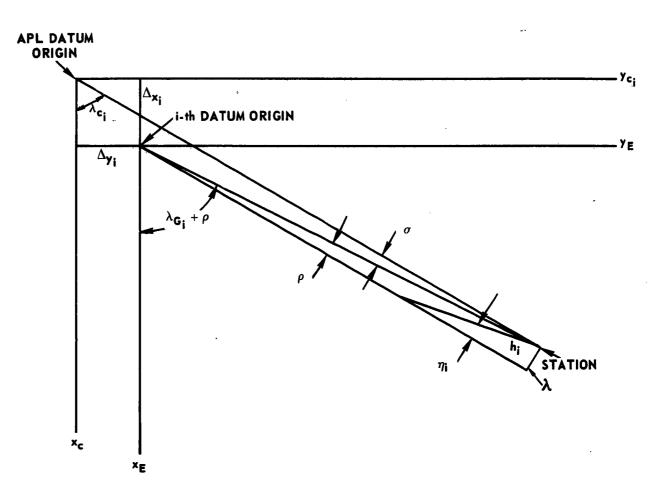


Figure 3B Diagram showing determination of λ_c when $h_i \, \not= \, 0.$

and it follows similarly

$$\rho \doteq \frac{\tau}{\zeta_{G_i}} = \frac{h_i \eta_i}{\zeta_{G_i}}$$

From equation (15) and Figure 3B,

$$\lambda_{c_{i}} = \lambda_{G_{i}} + \rho - \sigma$$

$$= \lambda_{G_{\underline{i}}} + \frac{h_{\underline{i}} \eta_{\underline{i}}}{\zeta_{G_{\underline{i}}}} - \frac{1}{\zeta_{G_{\underline{i}}}} [\Delta x_{\underline{i}} \sin (\lambda_{G_{\underline{i}}} + \rho) - \Delta y_{\underline{i}} \cos (\lambda_{G_{\underline{i}}} + \rho)]$$

Assuming $\cos \rho = 1 - \frac{\rho^2}{2}$, $\sin \rho = \rho$,

$$\lambda_{c_{i}} = \lambda_{G_{i}} + \frac{1}{\zeta_{G_{i}}} \left[h_{i} \eta_{i} - (\Delta x_{i} \sin \lambda_{G_{i}} - \Delta y_{i} \cos \lambda_{G_{i}}) \right] + 3rd \text{ order.}$$
 (16)

Equations (13), (14), and (16) are the cylindrical coordinates z_{c_i} , λ_{c_i} in the A.P.L. Earth fixed coordinate system expressed as a function of the geodetic coordinates of a tracking station.

References

- 1. Bomford, Brigadier G., "Geodesy", Clarendon Press, 1952.
- 2. Hasner, George L., "Geodesy", Wiley, Second Edit., 1930, (Chap. V Properties of the Spheroid and Chap. VIII Figure of the Earth).

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